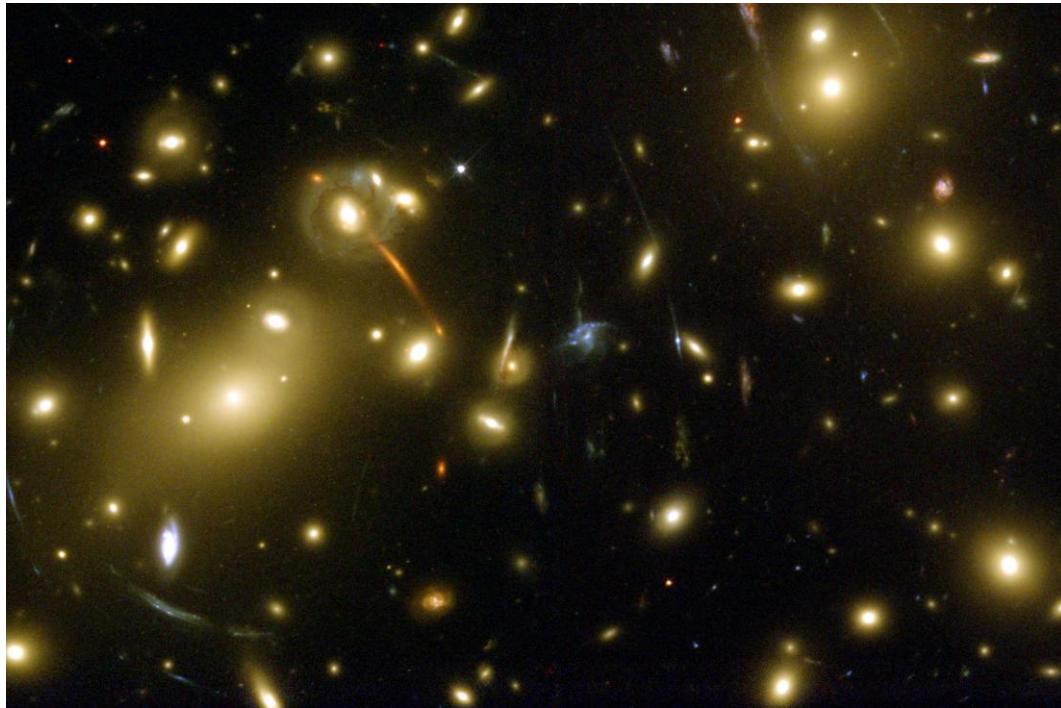


# The Role of Dark Matter in the Large-Scale Structures of the Universe



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# Agenda

- Introduction
- Types of matter
- Dark matter
  - Types of Dark Matter
  - Hot Dark Matter
  - Cold Dark Matter
  - Cold vs Hot Dark Matter
  - Cold and Hot Dark Matter
  - Cold Dark Matter and Dark Energy
- More models and theories
- Conclusions

# Introduction

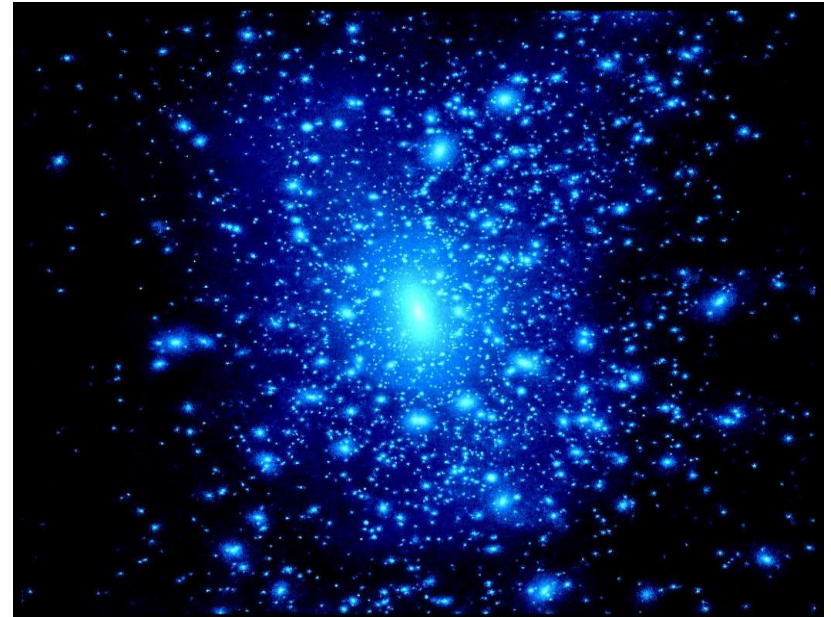
- Evidences of missing matter. Can Dark Matter be the answer?
- Concept introduced in the 30's but accepted until 70-80's
- Studied from Cosmological point of view
- Gap between cosmology and galaxy studies

# Types of matter

- Baryonic
  - Matter that forms atoms
  - Can contribute to the missing matter: black holes, neutron stars, white dwarfs, ...
- Non-baryonic
  - No emits, absorbs, scatter radiation
  - Gravitational interaction with normal matter

# Dark matter

- What's Dark Matter?
  - Nobody really knows
  - No emits, absorbs or scatter radiation
  - About 95% matter of the universe



# Dark Matter

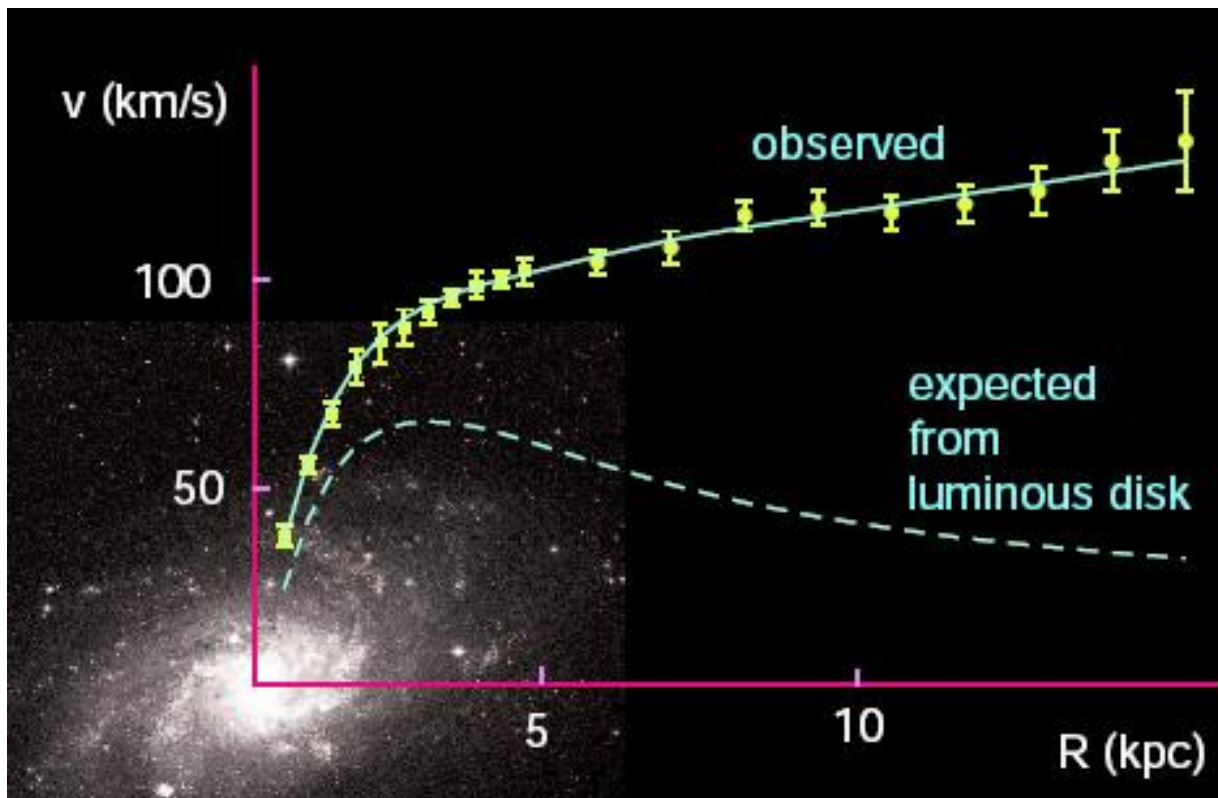
- Idea proposed in the 30's by Fritz Zwicky.

Why?

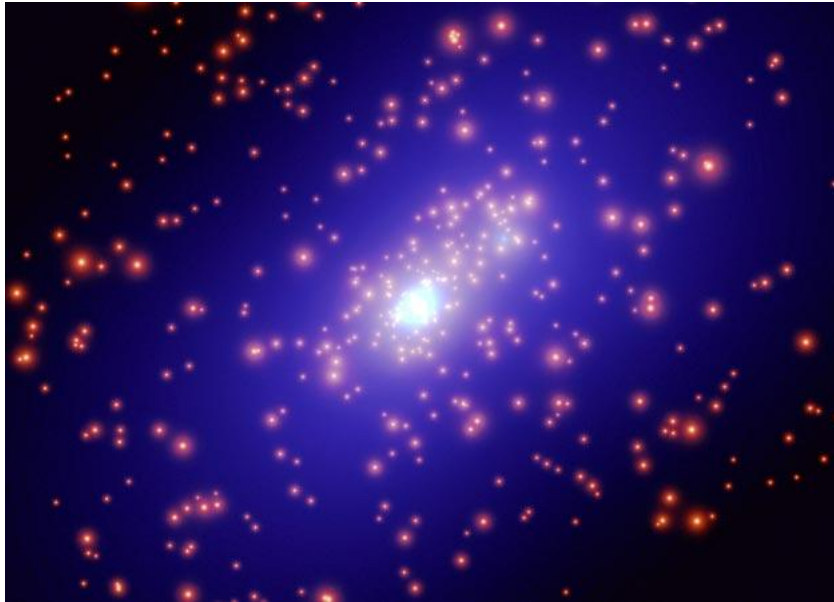
- Speed of Galaxies in large cluster too large
- Other evidences:
  - Rotation curves
  - Structure of galaxy groups and clusters
  - Etc.

# Dark Matter

## Example of rotation curve discrepancy



# Dark Matter: Types of Dark Matter

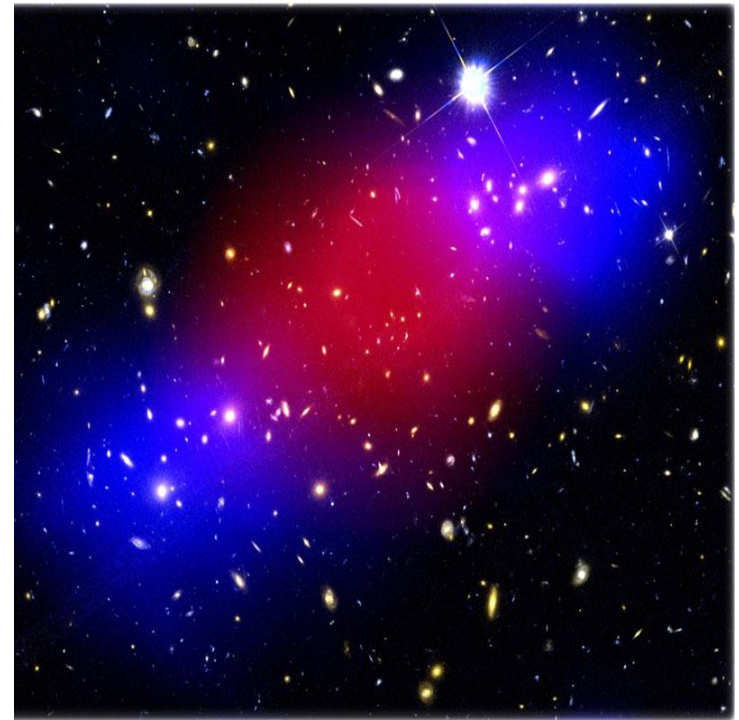


- Two main types
  - Hot Dark Matter (HDM)
  - Cold Dark Matter (CDM)



# Dark Matter: Hot Dark Matter

- Light neutrinos that decoupled when they had relativistic velocities
- Free streaming: Eliminate fluctuations in density smaller than superclusters  $\sim 10^{15} M_{\odot}$
- Top-down process: Fragmentation



# Dark Matter: Hot Dark Matter

- Good points
  - Structure of filaments is formed
- Bad points
  - Structures form too late
  - More inhomogeneous distribution than observed

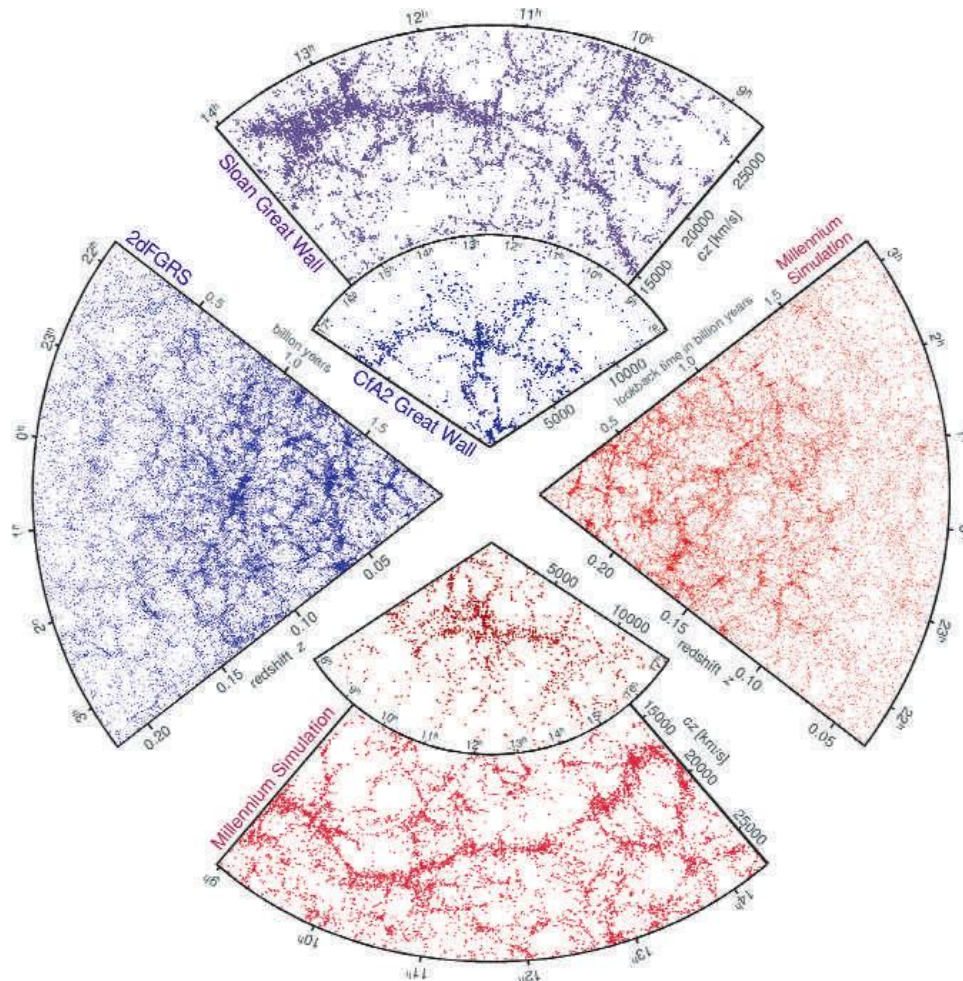
# Dark Matter: Cold Dark Matter

- Particles with Non-relativistic speed
- Possible candidates:
  - WMIP
  - Axions
- Perturbations at all scales survived, even small ones
- Bottom-up process: Hierarchical clustering

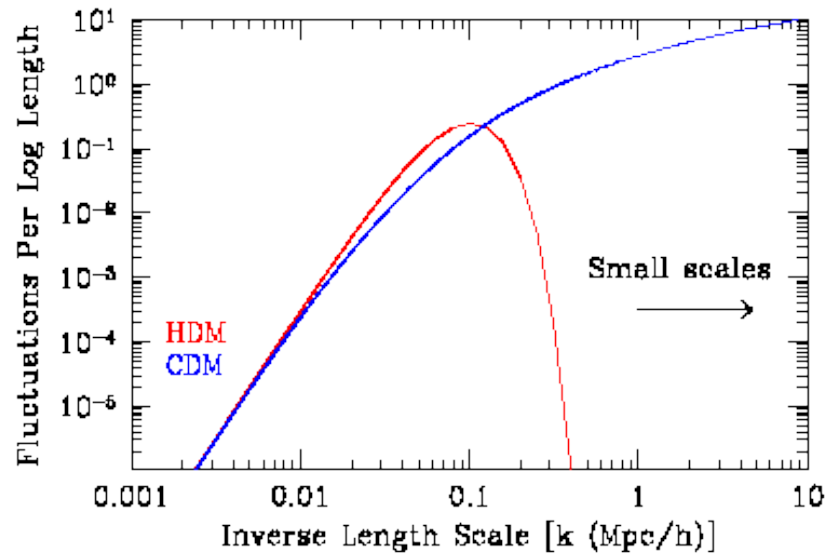
# Dark Matter: Cold Dark Matter

- Good points
  - Predicted distribution of galaxies agrees better with observations
- Bad Points
  - Missing satellite galaxies
  - Cuspy halo problem

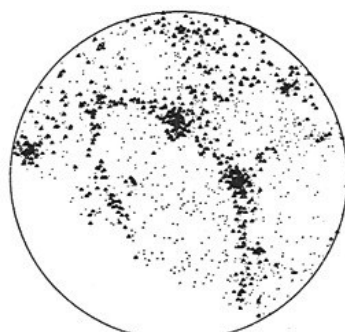
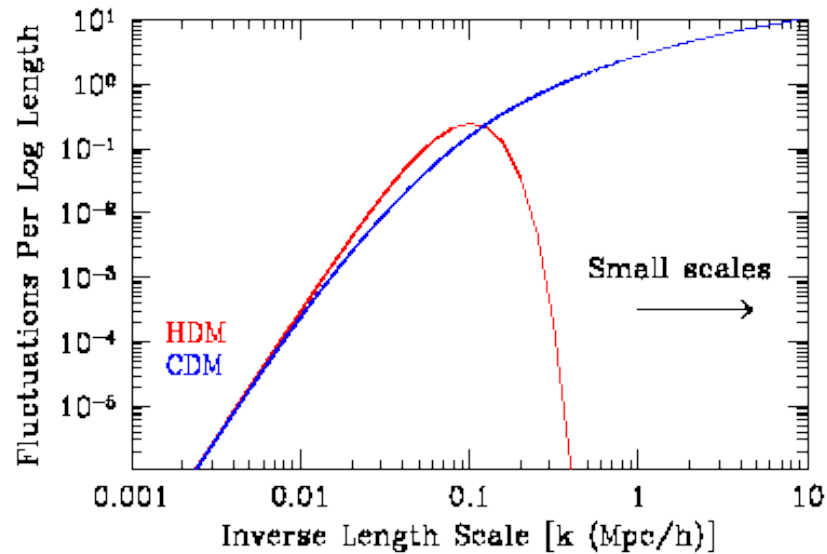
# Dark Matter: Cold Dark Matter



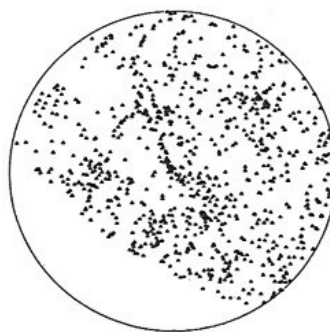
# Cold vs Hot Dark Matter



# Cold vs Hot Dark Matter



HDM



Observed Galaxy Distribution



CDM

# Dark Matter: Cold Hot Dark Matter

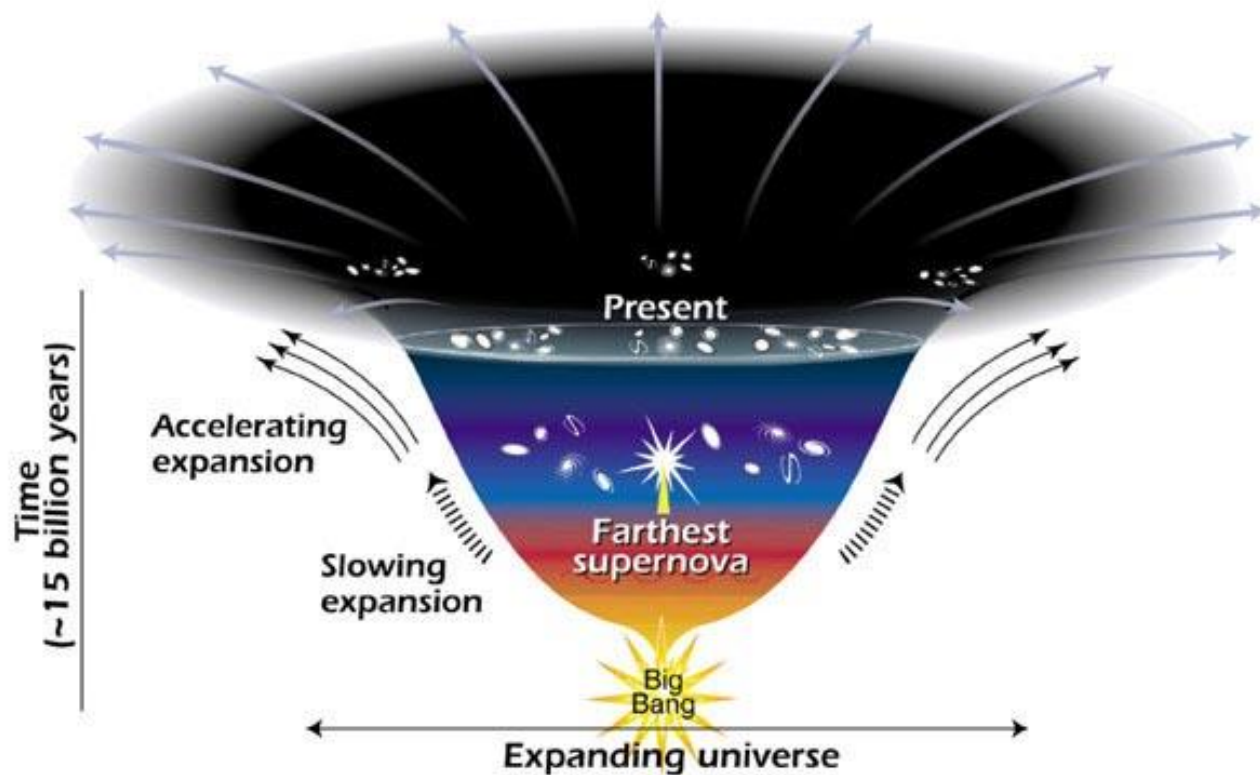
- Enough neutrino mass to take into account
- Small fluctuations can survive, just their grow rates are decreased
- About 20-30% of total matter is HDM
- Large matter density  $\Omega_m=1$



# Dark Matter: Cold Dark Matter and Dark Energy ( $\Lambda$ CDM)

- Inclusion of Dark Energy
- What is Dark Energy?
  - Some kind of negative pressure. Increase expansion rate of the universe
  - Evidences from Supernova Ia observations
- $\Omega_m \sim 1/3$     $\Omega_\Lambda \sim 2/3$     $\Omega_m + \Omega_\Lambda = 1$
- Observational data gives evidences favouring  $\Lambda$ CDM

# Dark Matter: Cold Dark Matter and Dark Energy ( $\Lambda$ CDM)



# More models and theories

- $\Lambda$ CHDM
  - Include neutrino mass (HDM) into the  $\Lambda$ CDM
- Modified Newtonian Dynamics (MOND)
  - Gravity depends on the scale of study

# Conclusions

- $\Lambda$ CDM is the most accepted model
- $\Lambda$ CHDM seems very promising
- Gap between Cosmological and smaller scale studies. Need to study this field

# Thank you!!

## Questions? : )

